

a)  $i\hbar \frac{d\hat{A}(t)}{dt} = i [\hat{A}(t), \hat{H}(t)] = 0 \Rightarrow \hat{A}(t) = \hat{A}$

b)  $\langle \alpha | \hat{A}(t) | \alpha \rangle = \langle \alpha | \hat{A} | \alpha \rangle = |\alpha|^2$

c) Consider  $\hat{H} |n\rangle = \left\{ n\hbar\omega_0 - \frac{\hbar k}{2} n(n-1) \right\} |n\rangle \Rightarrow$  states  $|n\rangle$  are eigenstates of  $\hat{H}$  with energy eigenvalues  $= n\hbar\omega_0 - \frac{\hbar k}{2} n(n-1)$

d)  $i\hbar \frac{d\hat{a}(t)}{dt} = [\hat{a}(t), \hat{H}(t)] = \hbar\omega_0 \hat{a}(t) - \hbar k \hat{a}^\dagger(t) \hat{a}(t) \hat{a}(t)$

$\Rightarrow \frac{d\hat{a}(t)}{dt} = -i\omega_0 \hat{a}(t) + ik \hat{a}^\dagger(t) \hat{a}(t) \hat{a}(t) = -i\omega_0 \hat{a}(t) + ik \hat{n}(t) \hat{a}(t)$   
 $= -i\omega_0 \hat{a}(t) + ik \hat{n} \hat{a}(t) = [-i\omega_0 + ik \hat{n}] \hat{a}(t)$

$\Rightarrow \hat{a}(t) = \exp[-i\omega_0 t + ikt \hat{n}] \hat{a}$

e) Need to find  $\langle \alpha | \hat{a}(t) | \alpha \rangle = \langle \alpha | e^{-i\omega_0 t} e^{ikt \hat{n}} \hat{a} | \alpha \rangle$

$= \alpha e^{-i\omega_0 t} \langle \alpha | e^{ikt \hat{n}} | \alpha \rangle = \alpha e^{-i\omega_0 t} \langle \alpha | \alpha e^{ikt} \rangle$

$= \alpha e^{-i\omega_0 t} e^{-|\alpha|^2 + |\alpha|^2 e^{ikt}} = \alpha e^{-i\omega_0 t} \exp[-|\alpha|^2(1 - \cos kt) + i|\alpha|^2 \sin kt]$

f) Energy lost from the cavity  $= [n\hbar\omega_0 - \frac{\hbar k}{2} n(n-1)] - [(n-1)\hbar\omega_0 - \frac{\hbar k}{2} (n-1)(n-2)]$

$= \hbar\omega_0 - \hbar k(n-1)$ . This energy must be taken away by the photon  $\Rightarrow$  freq. of photon that escaped  $= \omega_0 - k(n-1)$ . The spectrometer will measure this with probability 1.

g) A coherent state is a linear superposition of number states with probability  $\frac{|\alpha|^{2n}}{n!} e^{-|\alpha|^2}$  for number state  $|n\rangle$ . Spectrometer will measure frequency  $\omega_0 - k(n-1)$  with probability  $\frac{|\alpha|^{2n}}{n!} e^{-|\alpha|^2}$  for  $n=1, 2, 3, \dots, \infty$ .

h) we write a state as  $|n_1, n_2\rangle$ . Since  $n_1 + n_2 = 2$ , the three possible states are:  $|2, 0\rangle, |1, 1\rangle, |0, 2\rangle$ . Suppose the eigenvalue is  $E$ . We have  $\hat{H}|\psi\rangle = E|\psi\rangle$ . Taking the product of this equation with  $\langle 2, 0|$ , then  $\langle 1, 1|$ , and then  $\langle 0, 2|$  gives:

$$\begin{bmatrix} 2\hbar\omega_0 - \hbar k & -U & 0 \\ -U & 2\hbar\omega_0 & -U \\ 0 & -U & 2\hbar\omega_0 + \hbar k \end{bmatrix} \begin{bmatrix} a \\ b \\ c \end{bmatrix} = E \begin{bmatrix} a \\ b \\ c \end{bmatrix}$$

$$E_1 = 2\hbar\omega_0 - \hbar k$$

$$E_2 = 2\hbar\omega_0 - \frac{\hbar k}{2} - \frac{\sqrt{8U^2 + (\hbar k)^2}}{2}$$

$$E_3 = 2\hbar\omega_0 - \frac{\hbar k}{2} + \frac{\sqrt{8U^2 + (\hbar k)^2}}{2}$$

$$v_1 \propto \begin{bmatrix} 1 \\ 0 \\ 1 \end{bmatrix} \quad v_2 \propto \begin{bmatrix} 1 \\ -\frac{\hbar k}{2U} + \frac{\sqrt{8U^2 + (\hbar k)^2}}{2U} \\ 1 \end{bmatrix} \quad v_3 \propto \begin{bmatrix} 1 \\ -\frac{\hbar k}{2U} - \frac{\sqrt{8U^2 + (\hbar k)^2}}{2U} \\ 1 \end{bmatrix}$$

$$i) |\psi(t=0)\rangle = \frac{1}{\sqrt{2}} \left\{ |v_2\rangle - |v_3\rangle \right\} \Rightarrow \psi(t) = \frac{1}{\sqrt{2}} \left[ e^{-\frac{iE_2 t}{\hbar}} |v_2\rangle - e^{-\frac{iE_3 t}{\hbar}} |v_3\rangle \right]$$

2.

a) Since  $i(t) = f(t) + r(t) \Rightarrow n_i(t) = n_f(t) + n_r(t) = 0 \Rightarrow n_r(t) = -n_f(t)$ .

$$\Rightarrow \langle n_f(t+\tau) n_r(t) \rangle = -\langle n_f(t+\tau) n_f(t) \rangle = -\eta(1-\eta) \langle i(t) \rangle \delta(\tau)$$

b) The probability that  $n$  out of  $N$  particles incident on the splitter will get transmitted is  $= \frac{N!}{n!(N-n)!} \eta^n (1-\eta)^{N-n}$ . Therefore,

$$P_B(n, T) = \sum_{n=0}^{\infty} \frac{N!}{n!(N-n)!} \eta^n (1-\eta)^{N-n} \quad P_A(n, T) = \sum_{n=0}^{\infty} \frac{N!}{n!(N-n)!} \eta^n (1-\eta)^{N-n} \frac{(\alpha T)^n}{n!} e^{-\alpha T}$$

$$= \frac{(\eta \alpha)^n e^{-\eta \alpha T}}{n!}$$

3.

a) One can first boost the energy of level 1 to get the effective Hamiltonian:

$$\hat{H}_R = (\epsilon_3 - \Delta) |e_1\rangle \langle e_1| + \epsilon_3 |e_2\rangle \langle e_2| + \epsilon_3 |e_3\rangle \langle e_3| - \frac{\hbar \Omega_R}{2} \left[ e^{i\phi} |e_1\rangle \langle e_3| + e^{-i\phi} |e_3\rangle \langle e_1| \right] - U \left[ |e_2\rangle \langle e_3| + |e_3\rangle \langle e_2| \right]$$

Now this is equivalent to the EIT Hamiltonian of homework 4, problem 1.

Therefore:

$$\chi(\omega) = \frac{q^2 N (\vec{d}_{13} \cdot \hat{n})^2}{\epsilon_0 \hbar} \frac{-\Delta/\hbar}{i\gamma_{13} \frac{\Delta}{\hbar} - \left(\frac{\Delta}{\hbar}\right)^2 + \left(\frac{U}{\hbar}\right)^2}$$

b) when  $\Delta = 0$   $\text{Im}\{\chi(\omega)\} = 0$ .

c) The EIT condition is reqd:

$$\epsilon_3 - \epsilon_2 = \epsilon_3 - (\epsilon_1 + \hbar\omega) = \Delta$$